

The DUNE-PRISM Measurement Program for DUNE

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for the DUNE-PRISM Working Group

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DUNE Collaboration Call
June 22nd, 2018

Near Detector Group Recommendations

- ✦ Full report available in DUNE docdb 8184
- ✦ Recommendations excerpt:

R2) The design of a mobile LAr detector that can make measurements at one or more off-axis positions should go forward (DUNE-PRISM). The detector will need a side and downstream muon system for containment, which could either be the multi-purpose tracker (MPT) or a dedicated detector.

R3) Additional study of the DUNE-PRISM for technical feasibility and cost should be made.

R4) The underground experimental hall should be rotated by 90° in respect the the beam axis to allow for moving the near detector off axis.

- ✦ DUNE-PRISM technical note is also available in DUNE docdb 8106:

The DUNE-PRISM Near Detector Program

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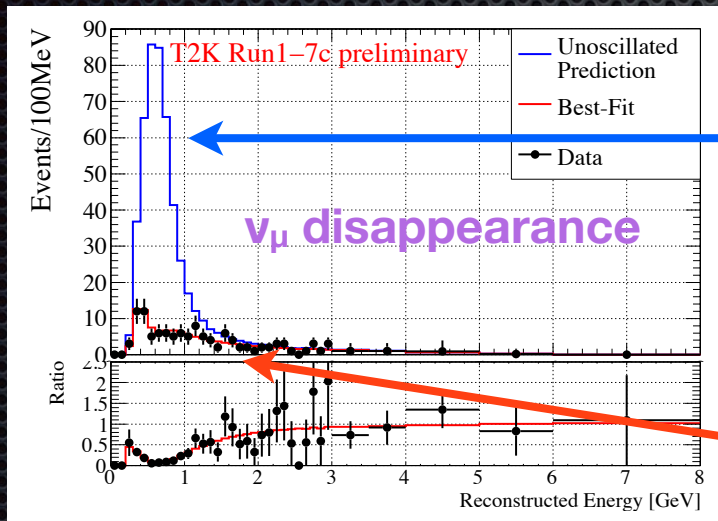
How to Measure Neutrino Oscillations

In a near/far experiment, σ uncertainties will cancel?

$$ND(\nu_\mu) = \Phi(E_\nu) \times \sigma(E_\nu, A) \times \epsilon_{ND} \times M_{E_{true}}^{E_{rec}}$$

$$FD(\nu_\mu) = \Phi(E_\nu) \times \sigma(E_\nu, A) \times \epsilon_{FD} \times P_{osc} \times M_{E_{true}}^{E_{rec}}$$

Cancellations of uncertainties in both flux and cross sections are spoiled by energy migrations

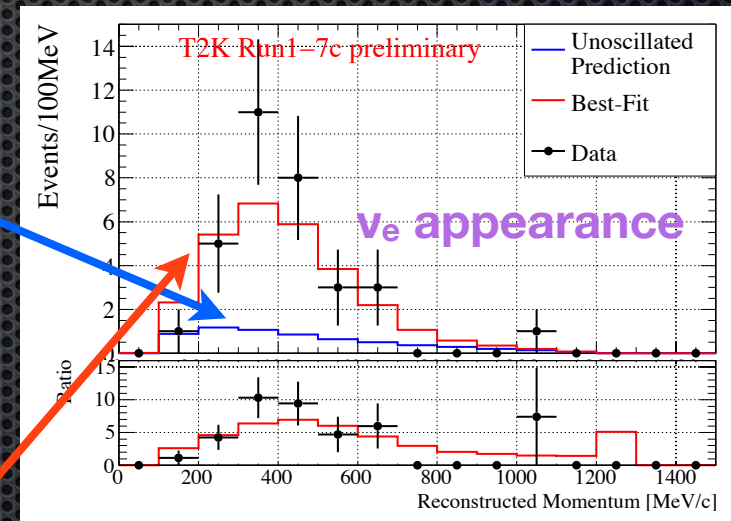


Near Detector Measures:

- ν_μ energy spectrum
- Small ν_e component

Far Detector Measures:

- Osc. ν_μ energy spectrum
- Large ν_e appearance signal

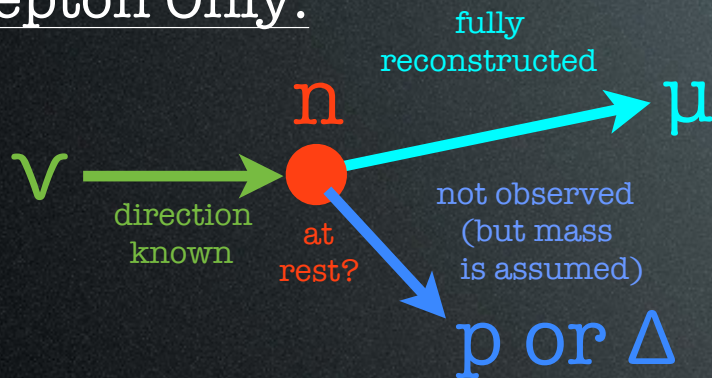


- ✦ $E_{true} \rightarrow E_{rec}$ migration matrix has significant off-axis components
 - ✦ Several important cross section uncertainties will not cancel

Measuring E_ν

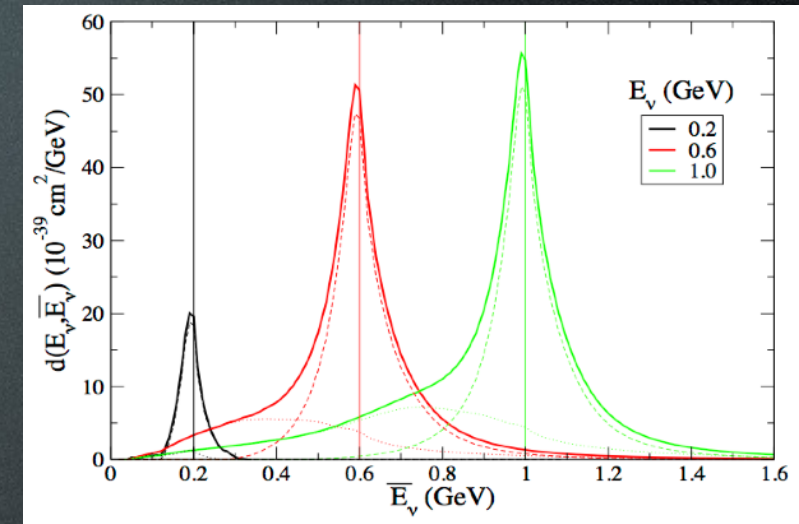
Martini et al. arXiv:
1211.1523

Lepton Only:

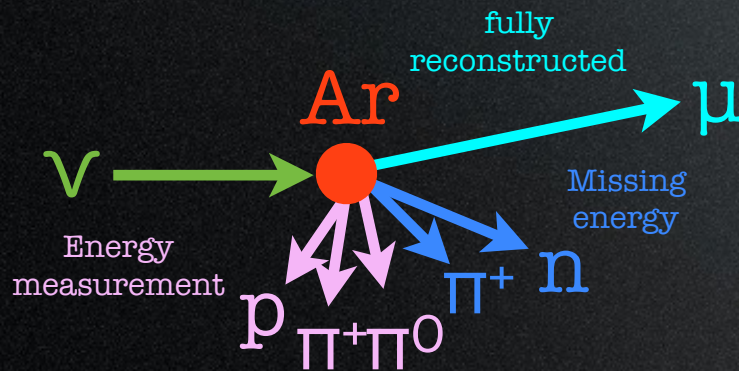


**Must assume
mass of
recoiling
hadron(s)**

**Problematic
due to
Multi-nucleon
interactions**



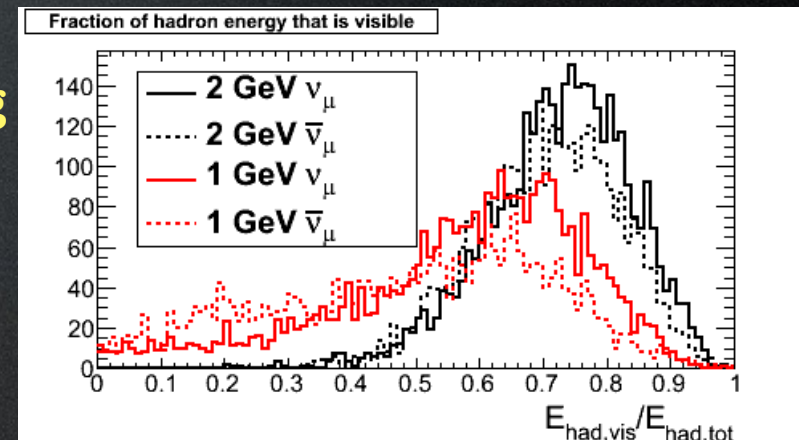
Lepton + Hadronic Energy:



**Missing hadronic
energy from n,
unseen π^+ , binding
energy, etc.**

**Energy loss
is different for
 ν and anti- ν**

[http://public.lanl.gov/friedland/LBNEApril2014/
LBNEApril2014talks/McGrew_LANL_Apr2014.pdf](http://public.lanl.gov/friedland/LBNEApril2014/LBNEApril2014talks/McGrew_LANL_Apr2014.pdf)



GEANT4 Simulation of a large LAr volume

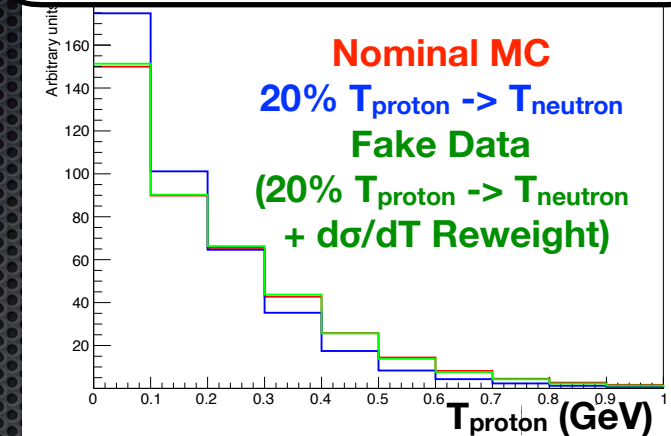
(True deposited hadronic energy)/
(True initial hadronic energy)

- Both effects lead to underestimating the neutrino energy (feed down)
- Need to calibrate both leptonic (e & μ) & hadronic energy scales and energy tails (variance)

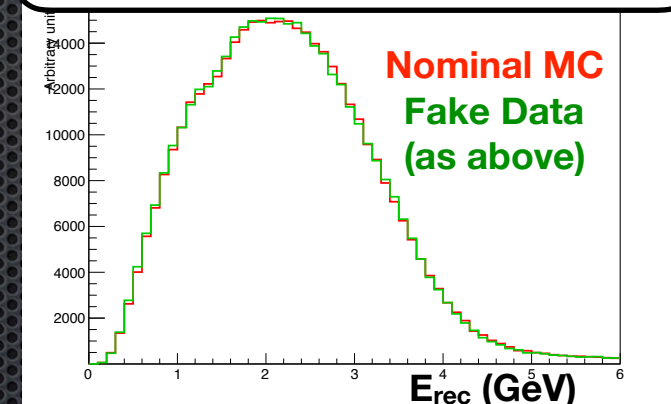
Fake Data Studies

- Suppose that 20% of the energy assumed to be emitted by protons is actually emitted by neutrons (which are unobserved)
- Experimentalists might correct the resulting data/MC discrepancy by reweighing the $d\sigma/dT$ distribution in the cross section model
 - This would be a typical result of fitting near detector data
- Both the T_{proton} and E_{rec} distributions can then agree perfectly
 - It looks like we perfectly understand our cross section model!
 - However, E_{rec} feed-down in data is much different than our model tells us

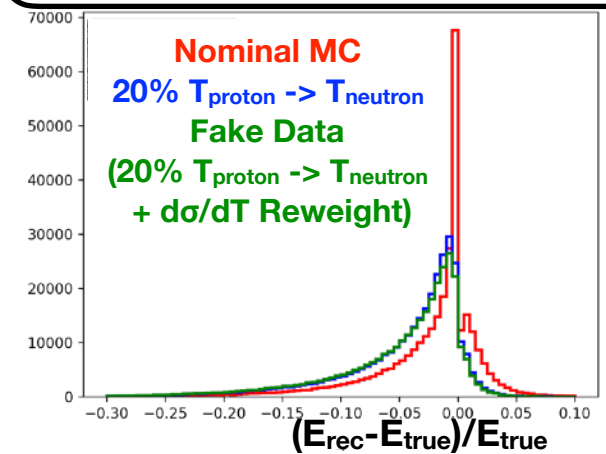
Near Detector T_{proton} On-Axis



Near Detector E_{rec} On-Axis

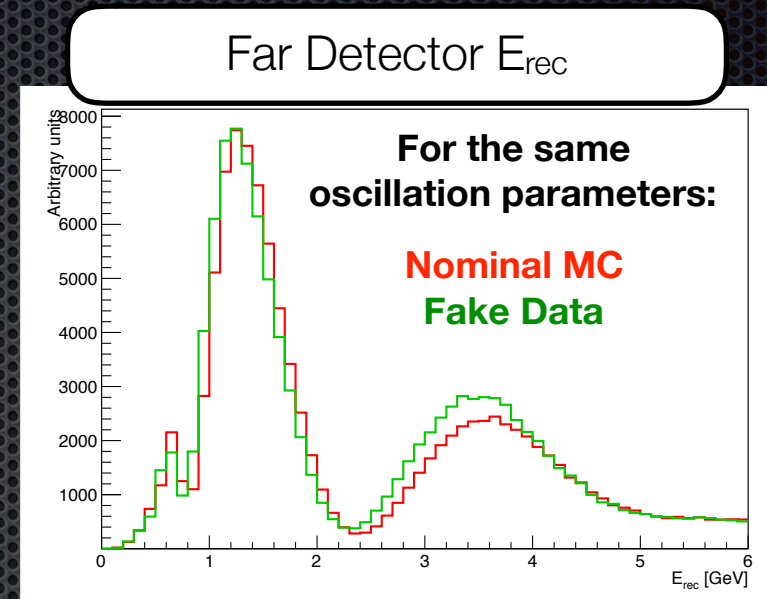


$E_{\text{true}} \rightarrow E_{\text{rec}}$

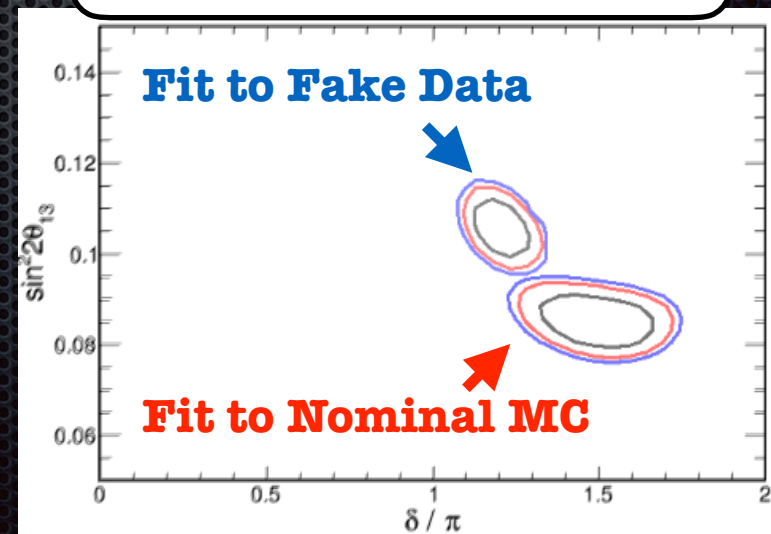


Fake Data II (Or: How to get the Wrong Answer for δ_{CP})

- Despite perfect agreement at the near detector, the fake data oscillation pattern at the far detector is wrong
 - $E_{true} \rightarrow E_{rec}$ is different than our model tells us
 - Even though our model was “confirmed” by our perfect near detector fit!
 - The E_{rec} distribution is shifted (wrong Δm^2) with a different magnitude oscillation dip (wrong θ_{23})
- Similarly, the ν_e distributions are also wrong
 - ...and by different amounts for neutrinos and anti-neutrinos
 - Hence, DUNE would measure the wrong value for δ_{CP}

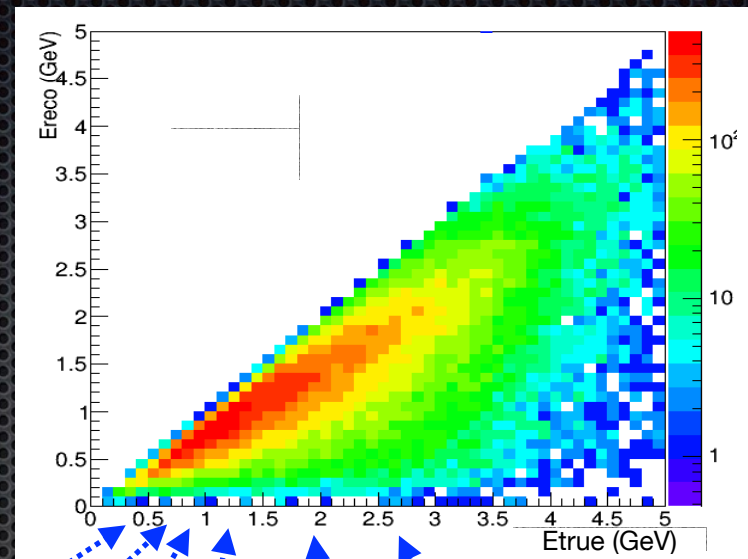


DUNE Oscillation Result
When Fitting the Fake Data or
the Nominal MC Prediction



DUNE-PRISM

- By moving the near detector to several off-axis positions, we can measure different E_ν spectra
- This provides a new degree of freedom over which we can constrain E_{rec} vs E_{true}
- There are various ways to combine such information to constrain the effects of cross section uncertainties on DUNE oscillation parameters
- (A few examples will be shown in this talk)



Beam

Increasing Off axis angle

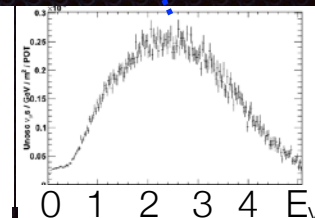
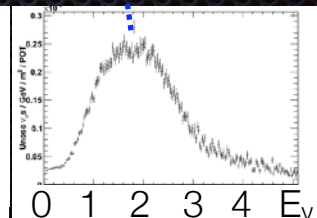
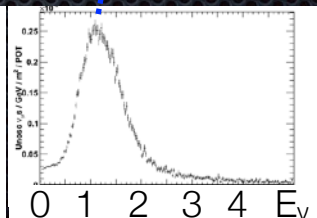
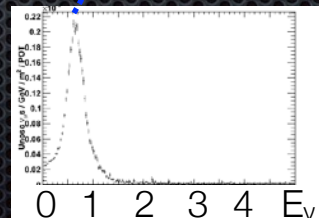
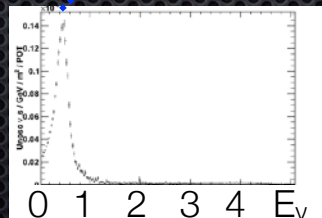
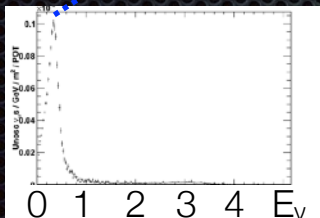
ND
Position
5

ND
Position
4

ND
Position
3

ND
Position
2

ND
Position
1

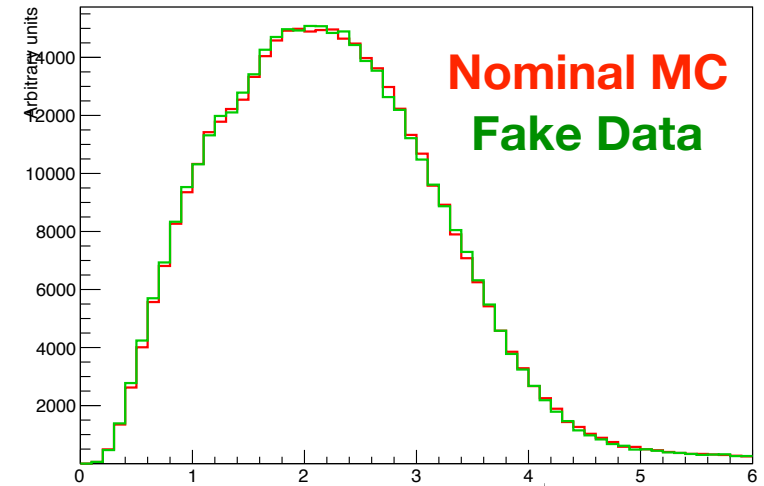


Fake Data w/ DUNE-PRISM

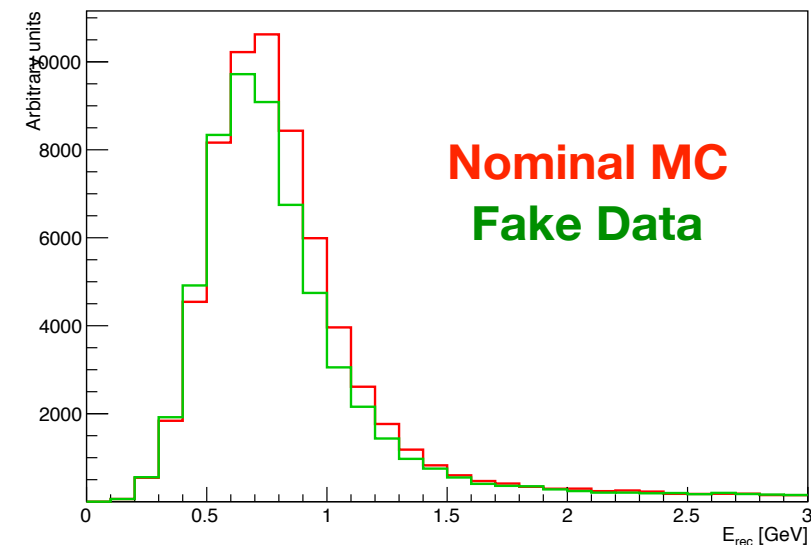
- By making a measurement in at least 1 off-axis location, formerly unseen problems in cross section modeling can be identified
 - e.g. we can avoid getting the wrong answer for e.g. δ_{CP}
- Now our cross section model $E_{true} \rightarrow E_{rec}$ matrix can be tested with 2 very different neutrino energy spectra
- We can do even better by making measurements at many off-axis locations
 - Particularly if we can continuously sample the whole off-axis range from ~33 m to on-axis

More details here: <https://indico.fnal.gov/event/14582/session/3/contribution/115/material/slides/>

Near Detector E_{reco} On-Axis

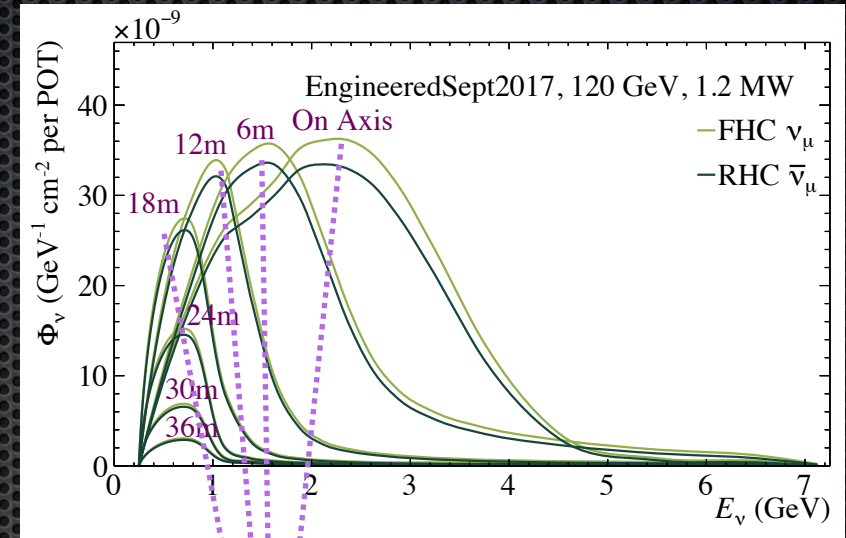


Near Detector E_{reco} 18 m Off-Axis

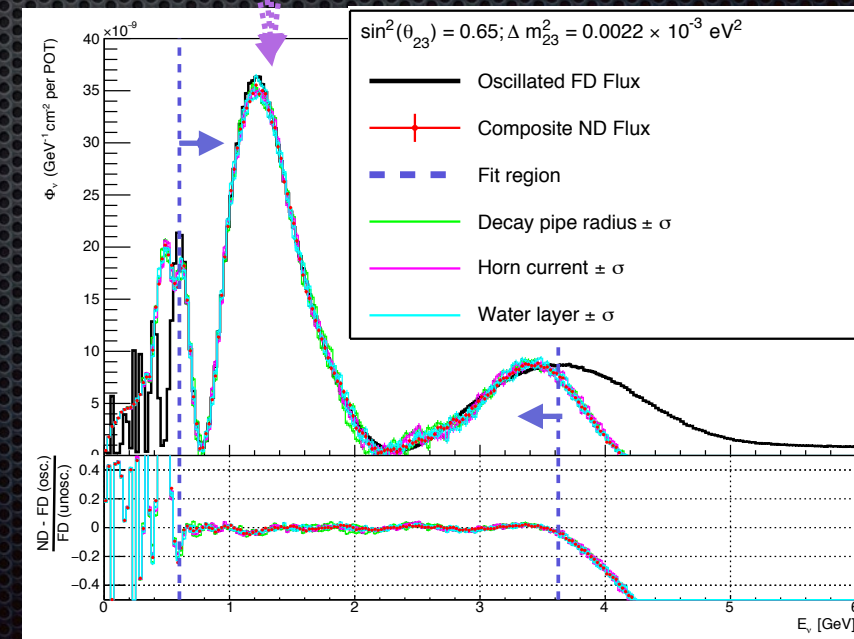


“Oscillated” Fluxes at the ND

- Recall: $E_{\text{true}} \rightarrow E_{\text{rec}}$ migrations are problematic due to near \rightarrow far flux differences (due to oscillations)
- We can construct at “oscillated” flux at the near detector using linear combinations of off-axis measurements
- These linear combinations are largely insensitive to the expected flux systematic uncertainties

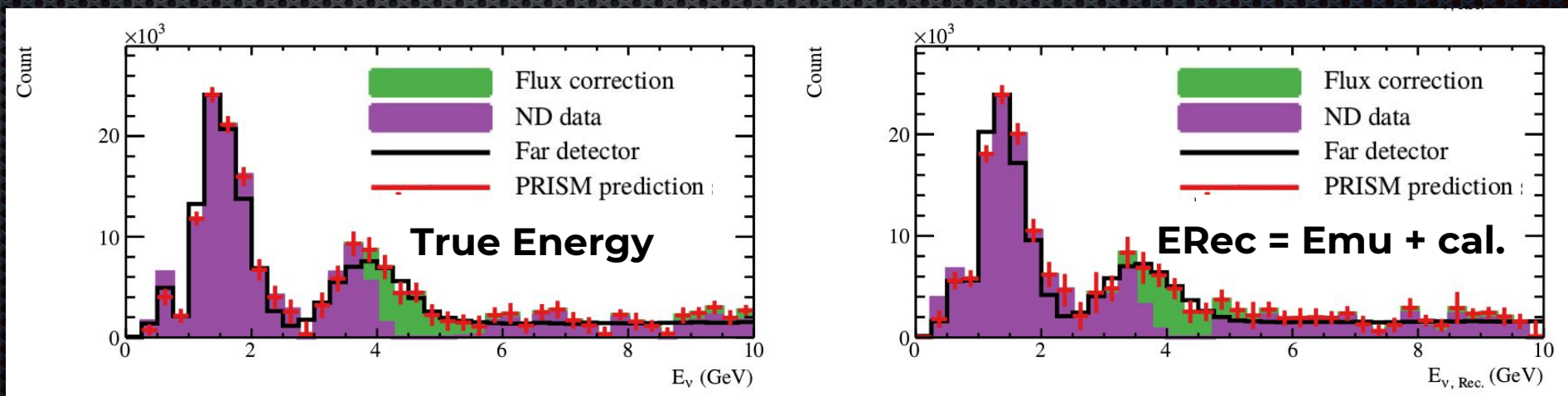


Linear Combinations of measurements at different off-axis angles



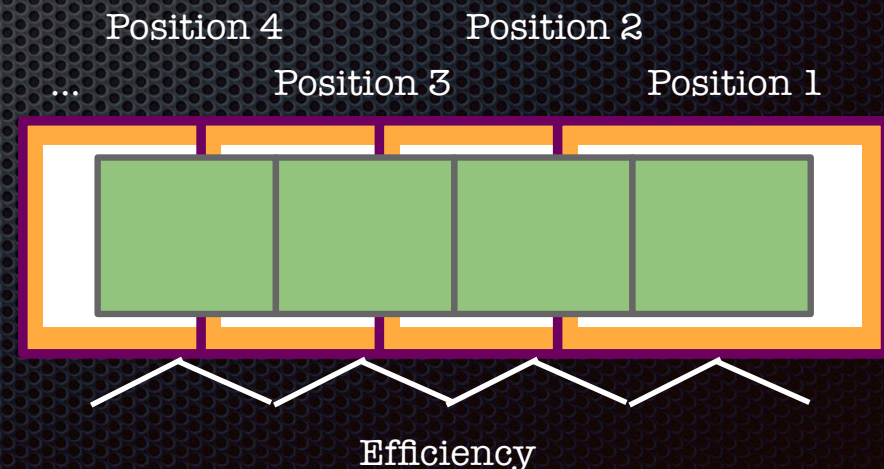
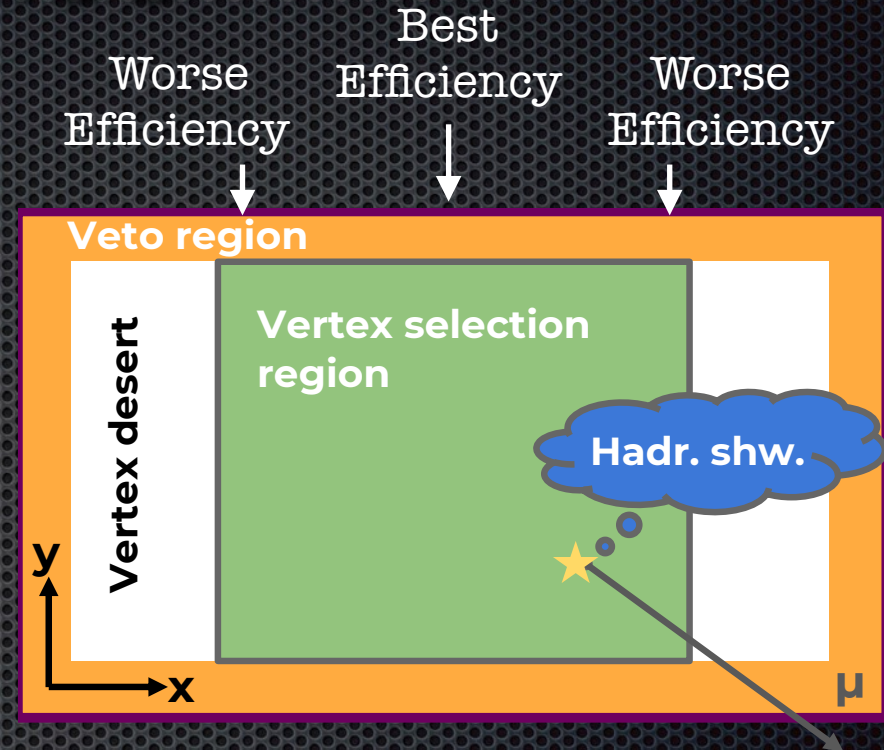
Data-Driven Oscillated E_{rec}

- ✦ The flux-fit linear combinations are applied to the measured E_{rec} distributions at each off-axis location
 - ✦ This gives a data-driven estimate of the E_{rec} distribution we would see at the far detector
 - ✦ Near to far extrapolation is now independent of GENIE to first order (Residual model dependence remains in background subtraction and flux fit corrections)
- ✦ Work-in-progress; efficiency correction still under study (next slide)



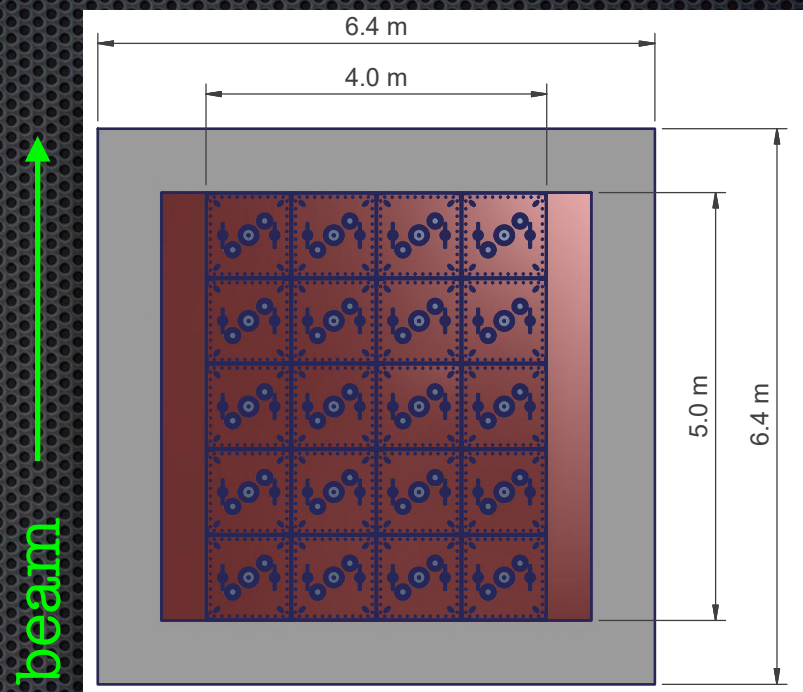
Off-Axis Efficiencies

- Ideally, we would like uniform efficiency across all off-axis positions
 - Current event selection requires < 20 MeV of hadronic energy in our 50 cm of LAr active region
- Unfortunately, hadronic showers can more easily reach the veto region when closer to the edge of the detector
- Fiducial volume is reduced in off-axis direction so selected events are less likely to trigger the hadronic shower veto
 - This makes the efficiency more uniform vs off-axis angle



ND LAr Detector Size & Movement

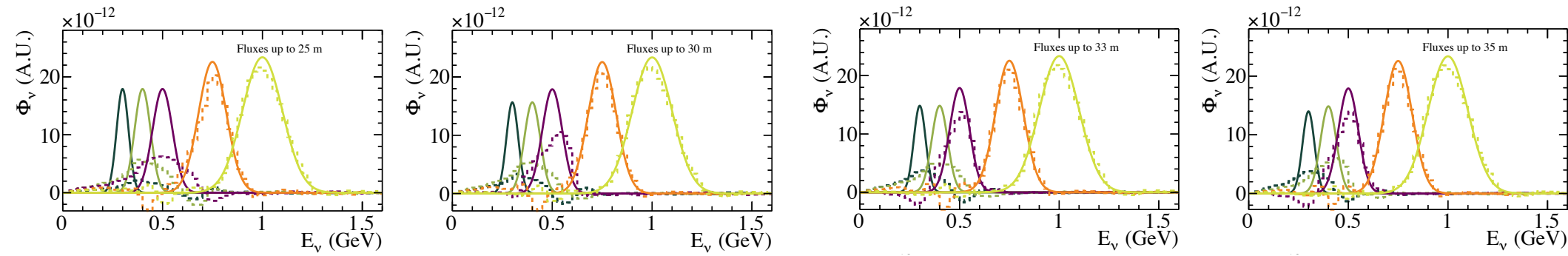
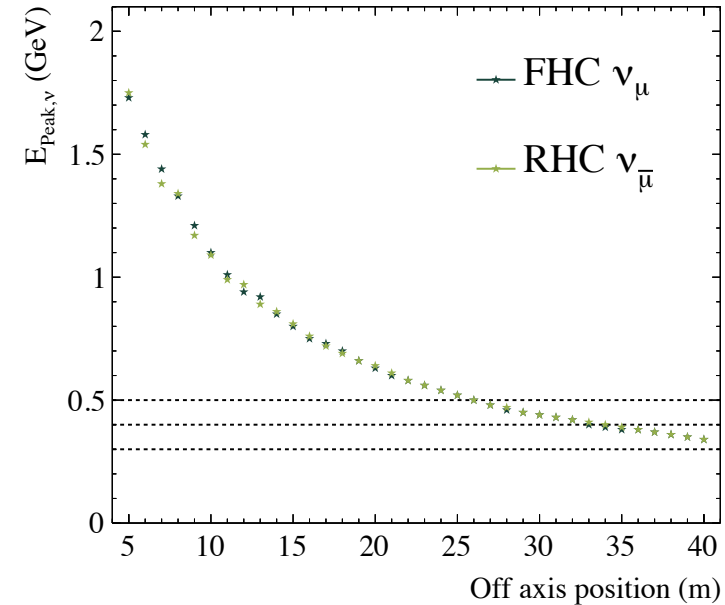
LAr Dimensions



- ✦ Originally assumed LAr size was 4 m wide (off-axis) x 3 m (tall) x 5 m (beam direction)
 - ✦ This is the “minimal” size for hadronic shower containment
 - ✦ However, to contain sideways-going muons, off-axis dimension should be increased (4 m → 7 or 8 m)
 - ✦ Removes the need for a side muon detector
 - ✦ For the DUNE-PRISM program, a wider detector means less measurement positions, and more uniform efficiency
- ✦ 3 options under investigation by Fermilab engineers:
 1. A continuously moving LAr detector
 2. Moving to fixed positions
 3. A fixed, ~35 m wide LAr detector

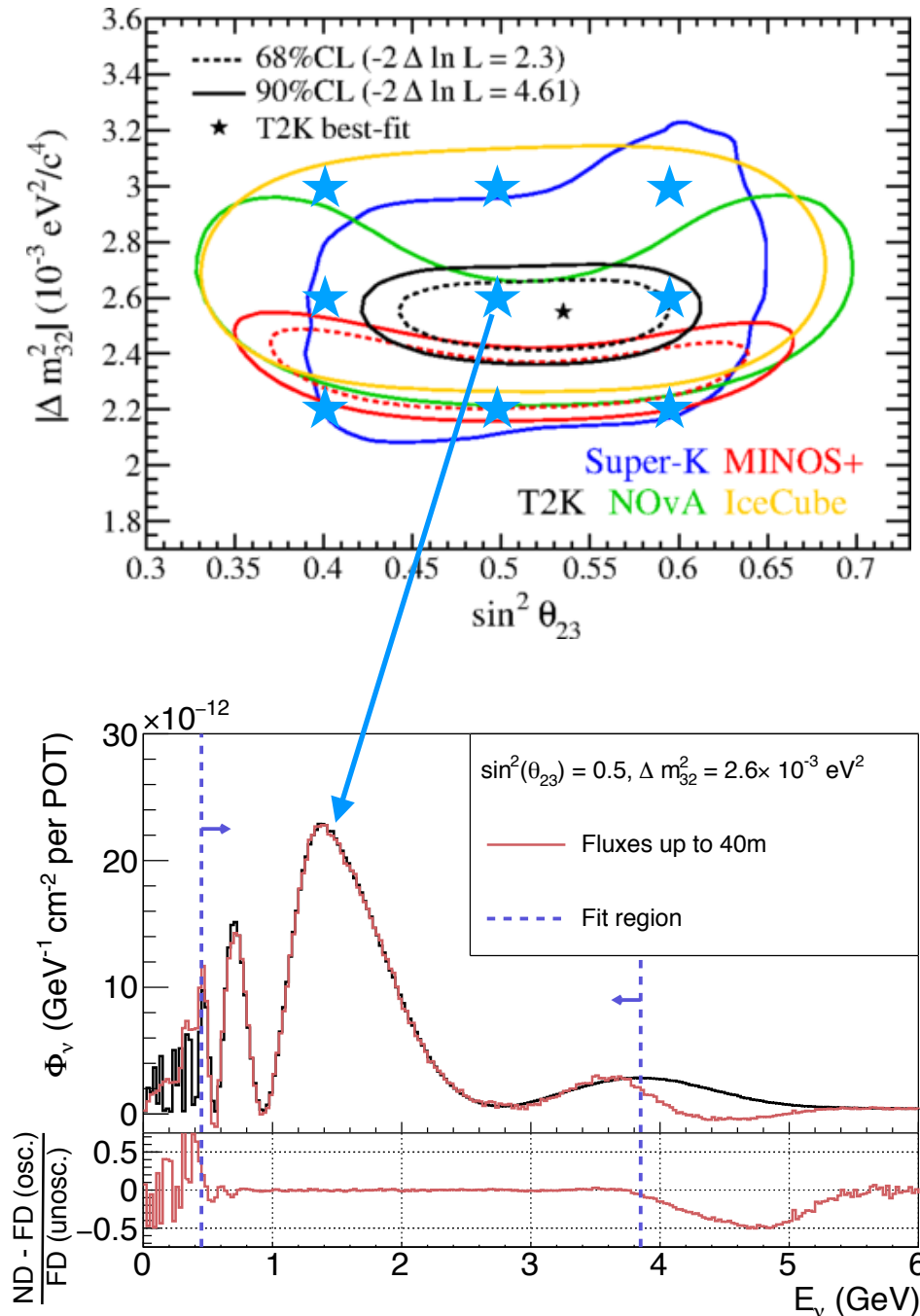
How Far Off-Axis?

- Further off-axis = lower reach in neutrino energy
- 500 MeV flux peaks at 26 m off-axis
 - To properly understand events at 500 MeV, we need access lower energies at further off-axis positions
- One method to determine the lowest needed energy is to construct a Gaussian energy spectrum at 500 MeV (10% width) using linear combinations of off-axis fluxes
 - This is not the only method one could employ (see next slides)
 - The 500 MeV Gaussian fit clearly begins to degrade when fluxes between 30 m & 33 m are excluded



Oscillated Flux Fits

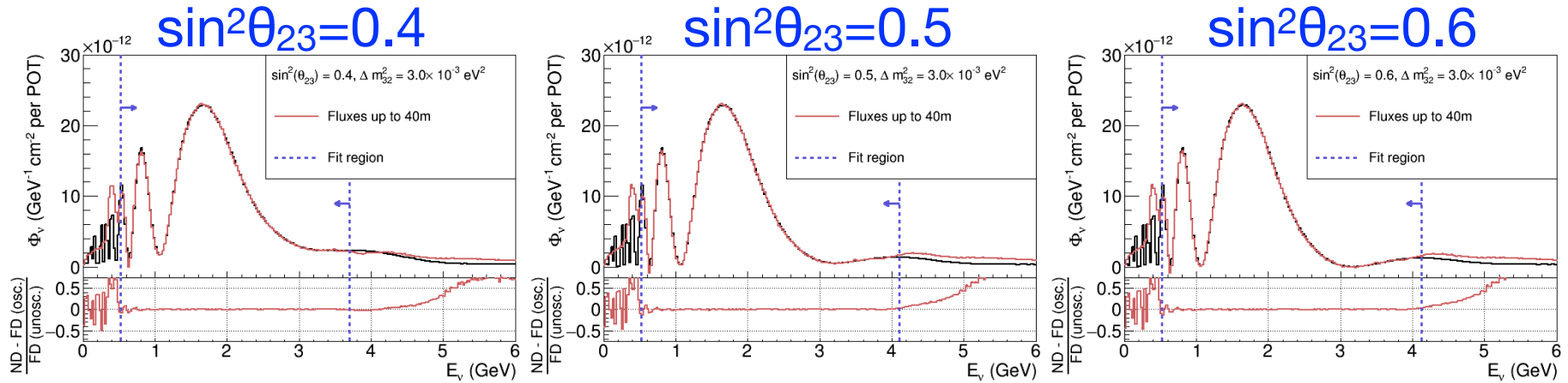
- We can also use linear combinations of off-axis fluxes to construct an oscillated flux seen at the far detector for any currently allowed set of oscillation parameters
- Again, this is not the definitive metric, but it does show how well such a fit can resolve the bump below the 2nd oscillation maximum (which peaks as low as ~ 500 MeV, depending on Δm_{32}^2)
- The following slides probe the 9 points in $\Delta m_{32}^2, \theta_{23}$ space shown in the top figure
 - Vary off-axis range used in fits



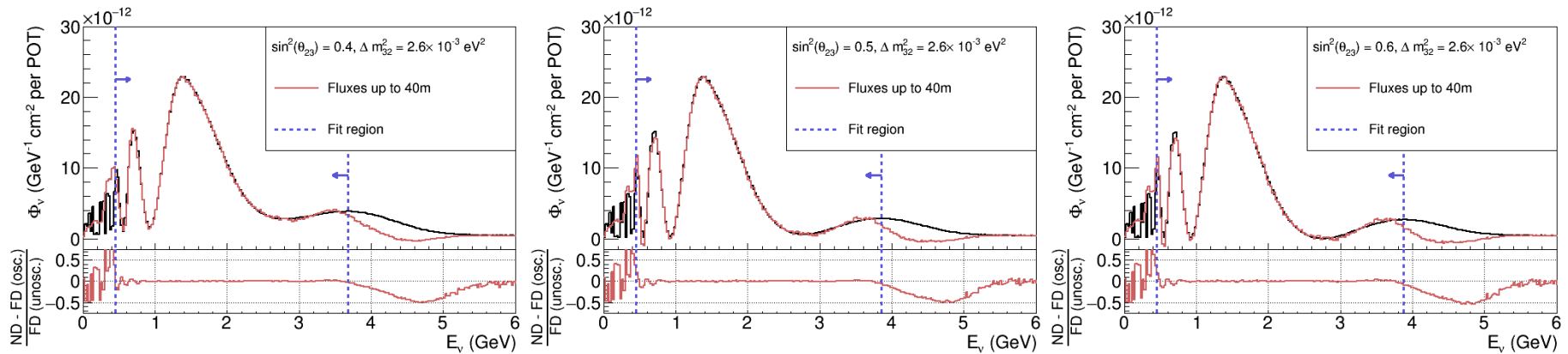
Fluxes Up to 40 m Off-Axis

- More off-axis range than needed. We can even somewhat resolve the peak below the 3rd oscillation maximum for all values of Δm_{32}^2

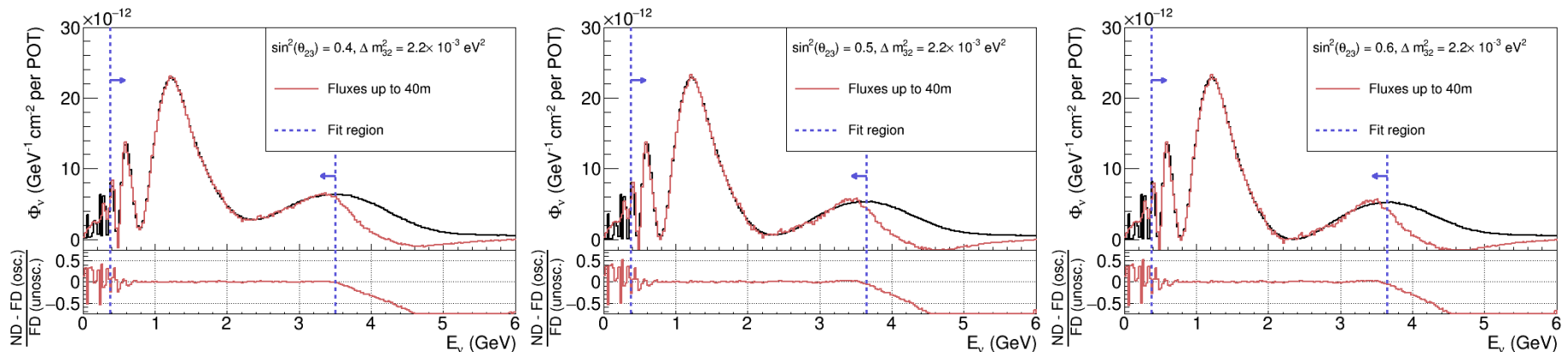
$\Delta m^2 = 3.0 \times 10^{-3}$



$\Delta m^2 = 2.6 \times 10^{-3}$



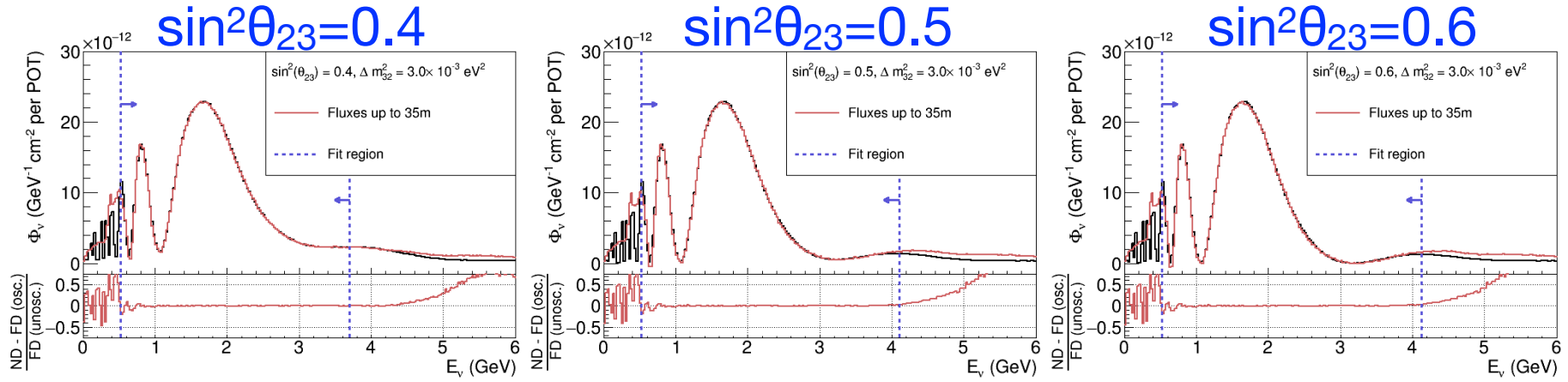
$\Delta m^2 = 2.2 \times 10^{-3}$



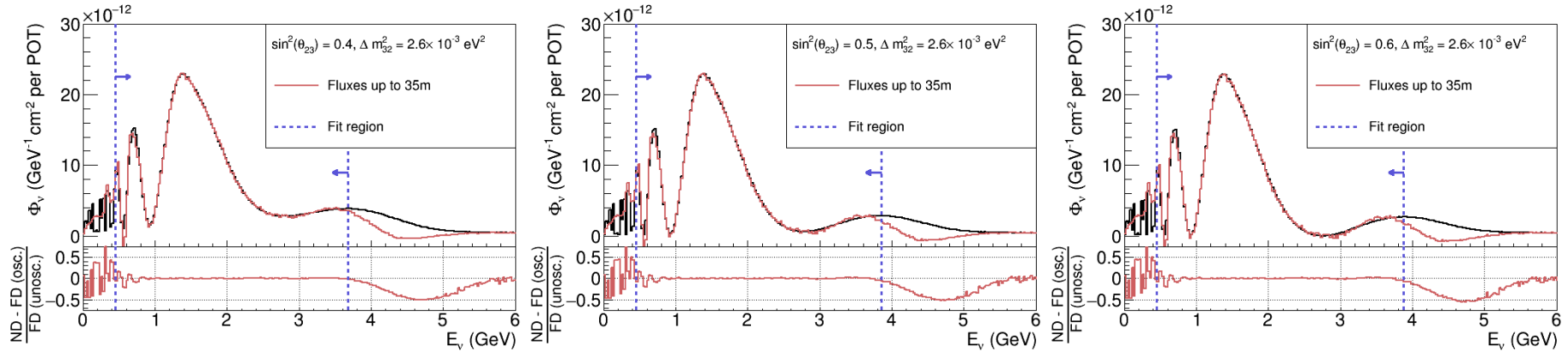
Fluxes Up to 35 m Off-Axis

- Can generally resolve bump below 2nd oscillation maximum for all values of Δm_{32}^2 , although some fluctuations are seen in the ratio to the unoscillated flux

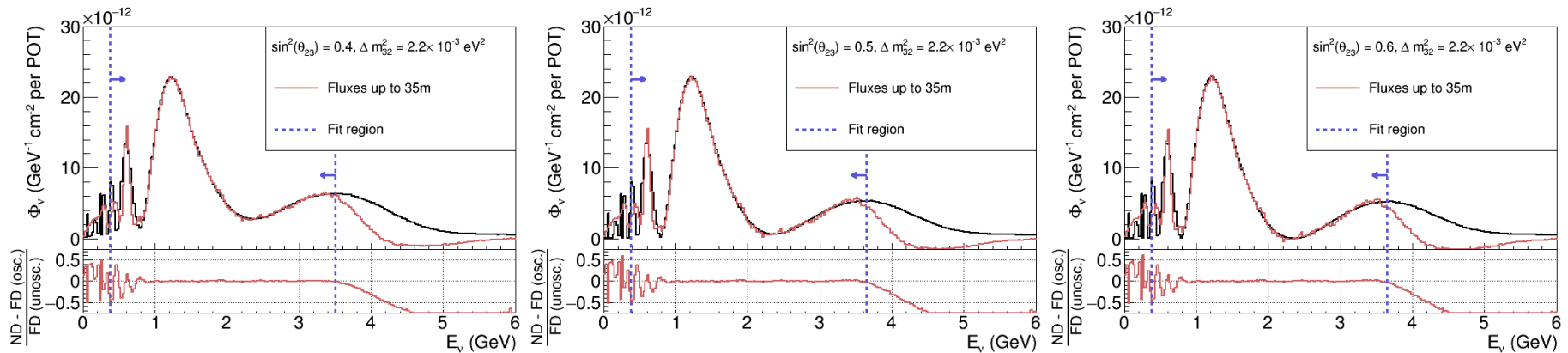
$\Delta m^2 = 3.0 \times 10^{-3}$



$\Delta m^2 = 2.6 \times 10^{-3}$



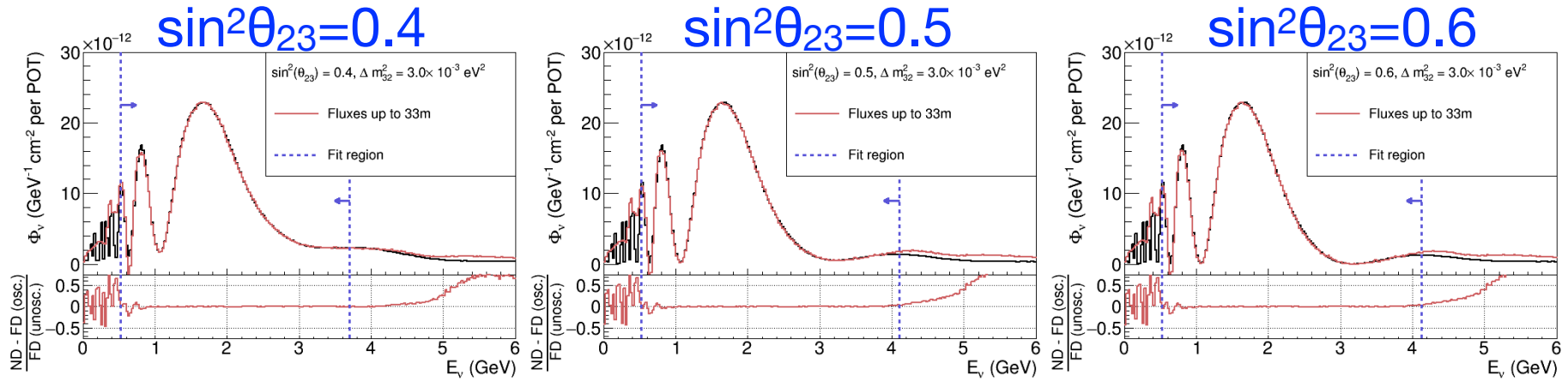
$\Delta m^2 = 2.2 \times 10^{-3}$



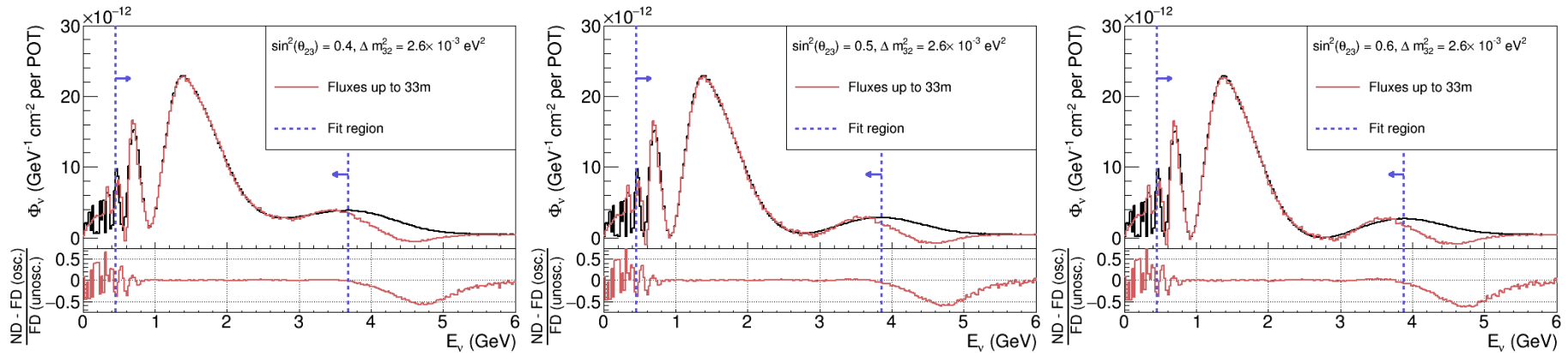
Fluxes Up to 33 m Off-Axis

- Can still generally resolve bump below 2nd oscillation maximum for all values of Δm_{32}^2 , although some fluctuations are seen in the ratio to the unoscillated flux

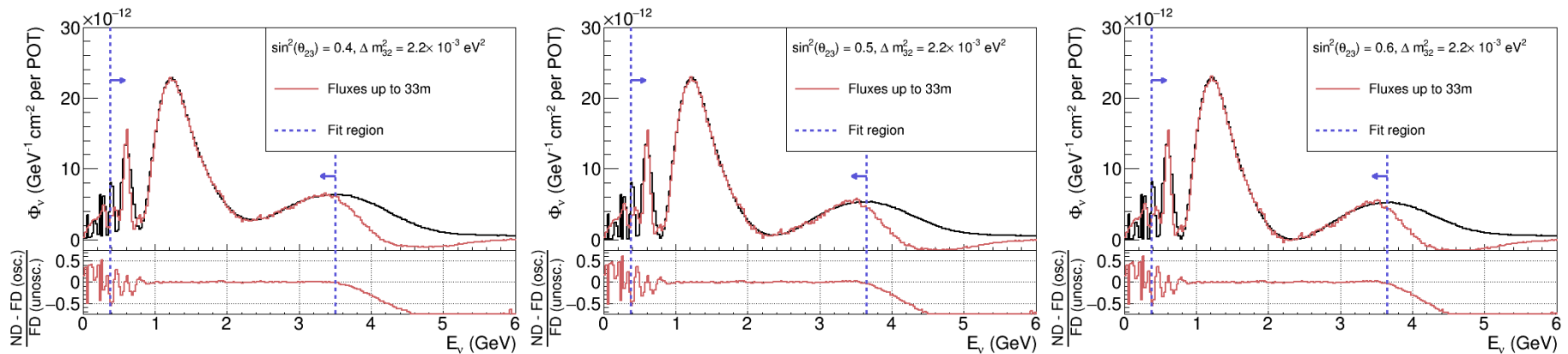
$\Delta m^2 = 3.0 \times 10^{-3}$



$\Delta m^2 = 2.6 \times 10^{-3}$



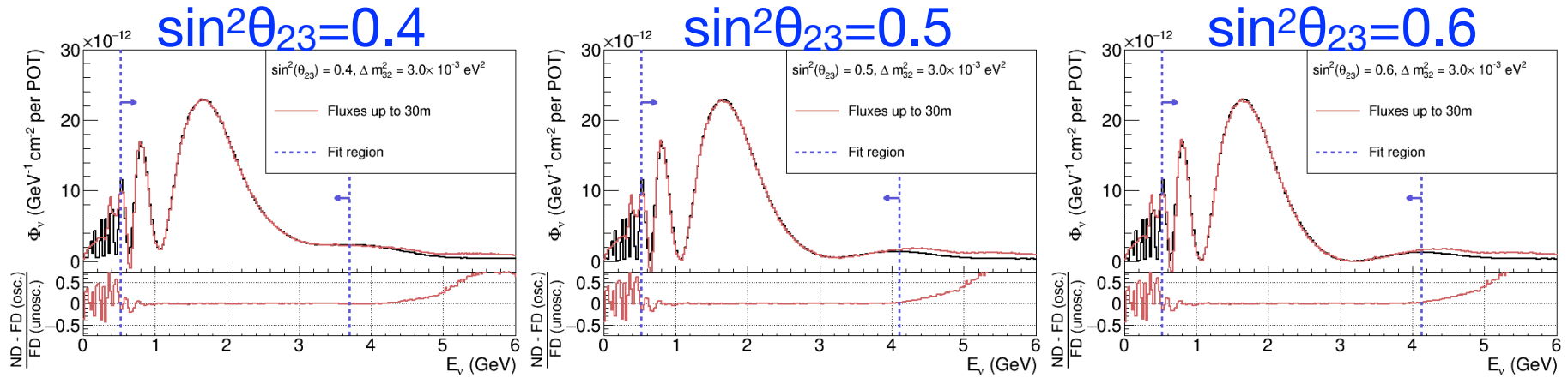
$\Delta m^2 = 2.2 \times 10^{-3}$



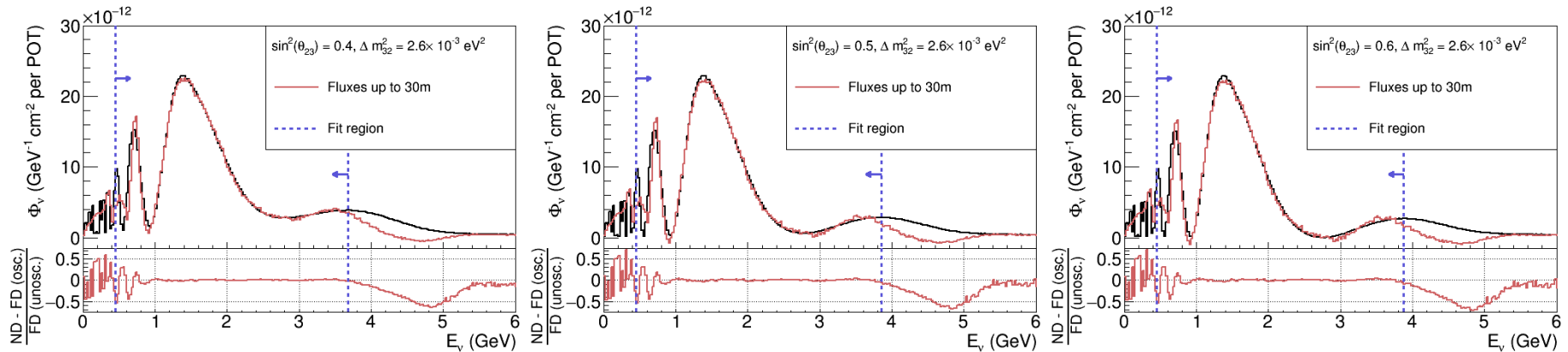
Fluxes Up to 30 m Off-Axis

- Poor fits around the 2nd oscillation maximum for low Δm_{32}^2 region; ability to constrain systematics in this region may be compromised

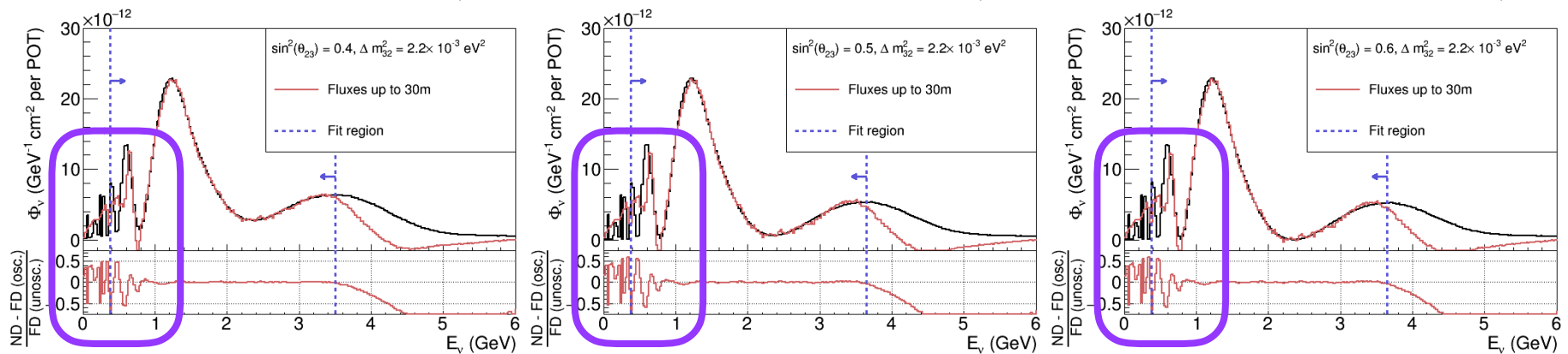
$\Delta m^2 = 3.0 \times 10^{-3}$



$\Delta m^2 = 2.6 \times 10^{-3}$



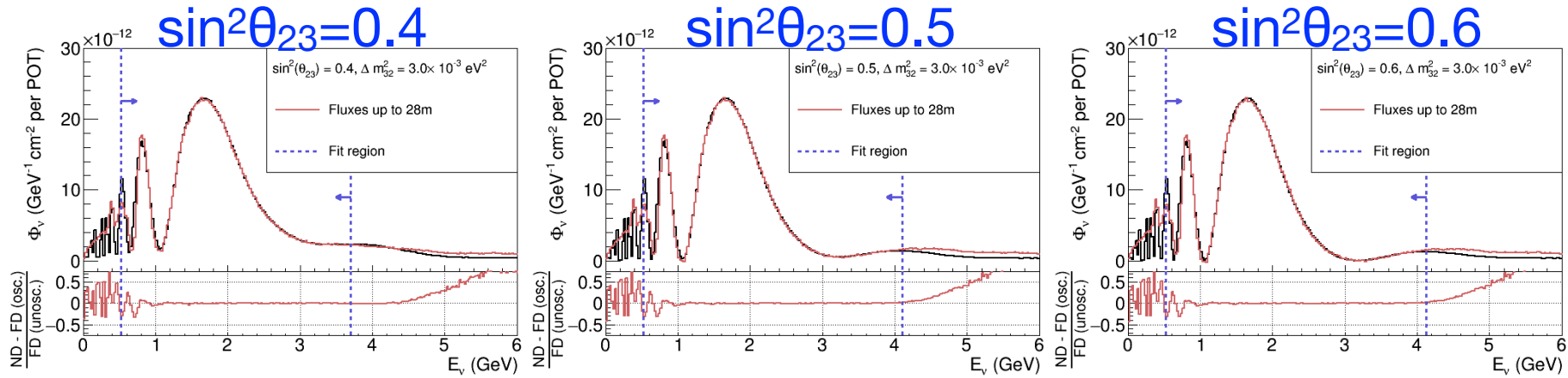
$\Delta m^2 = 2.2 \times 10^{-3}$



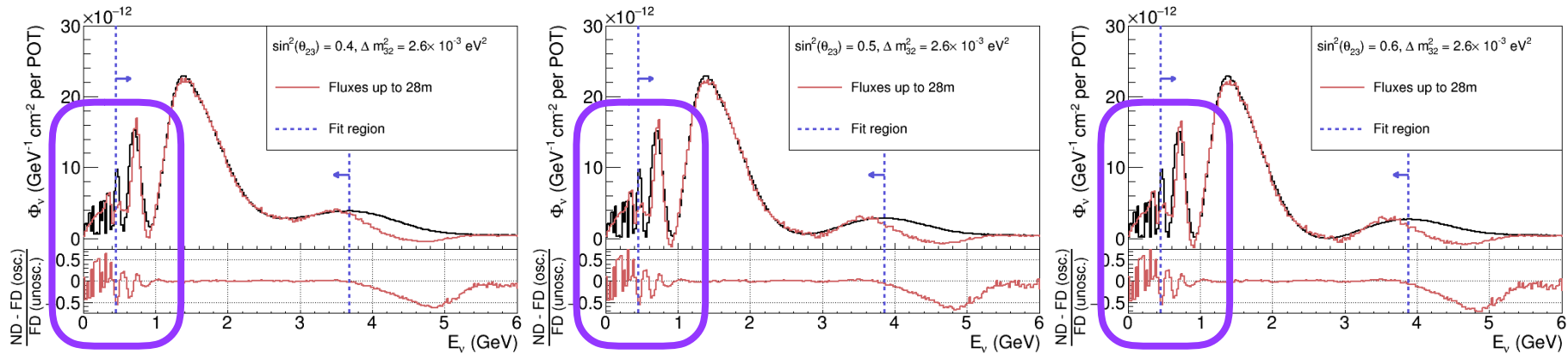
Fluxes Up to 28 m Off-Axis

- Very poor fits around the 2nd oscillation maximum for low Δm_{32}^2 ; limiting to 28 m can cause harm to 2nd oscillation maximum physics

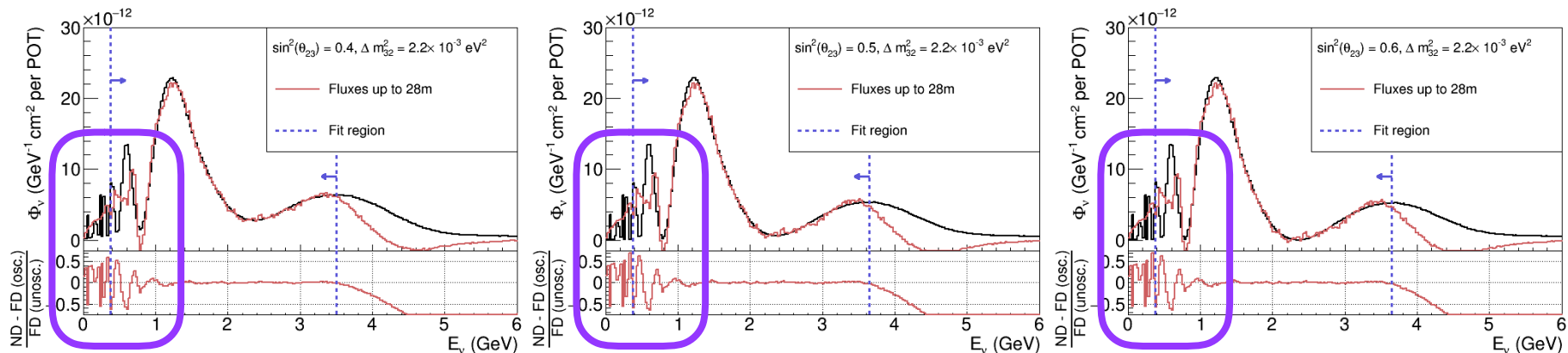
$\Delta m^2 = 3.0 \times 10^{-3}$



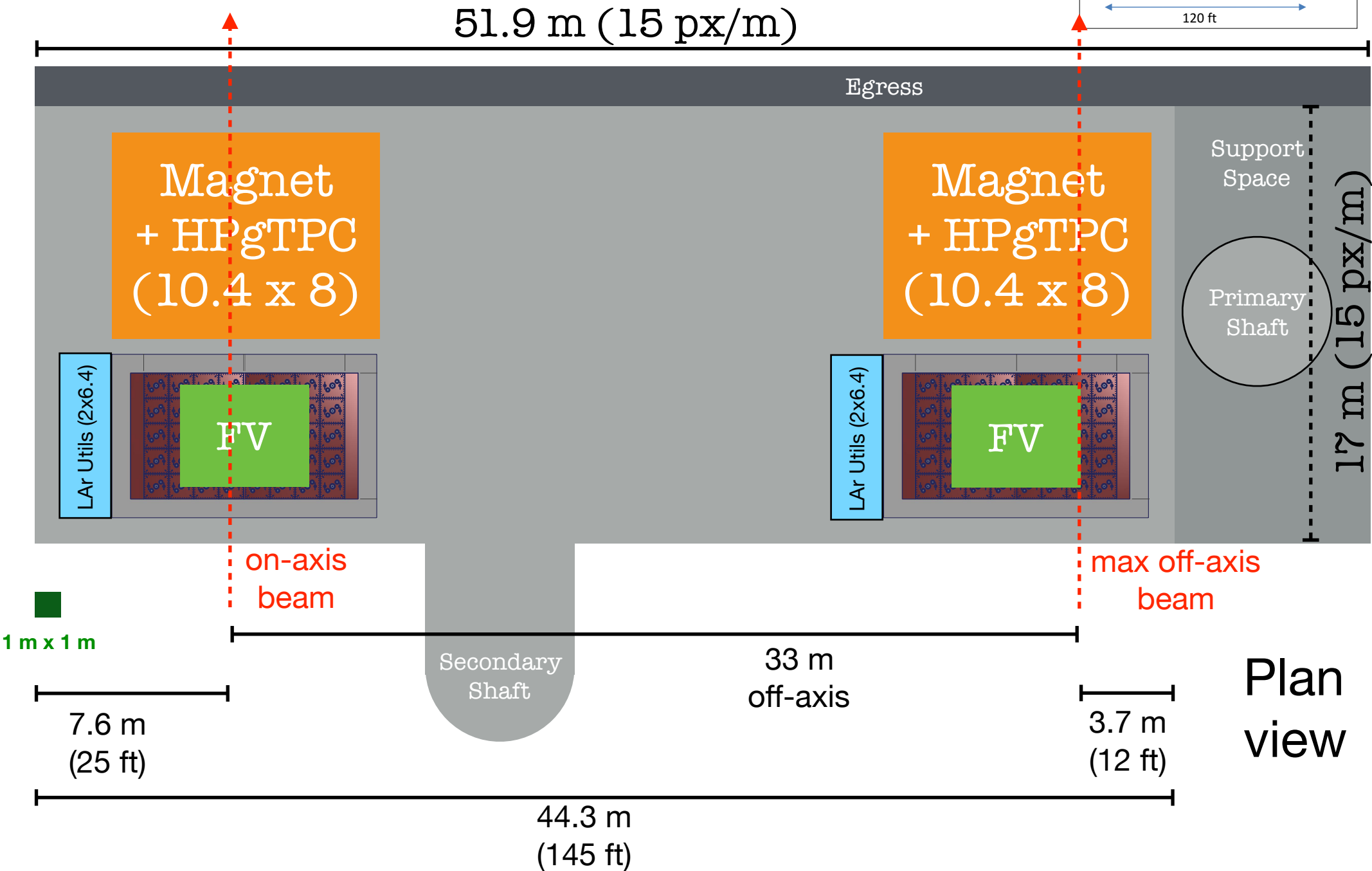
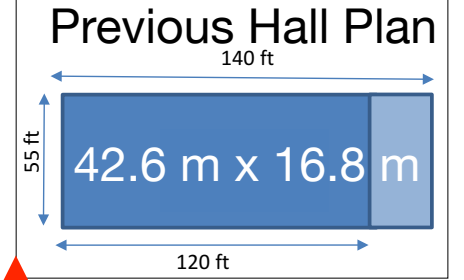
$\Delta m^2 = 2.6 \times 10^{-3}$



$\Delta m^2 = 2.2 \times 10^{-3}$



Ideal ND Hall Layout



Conclusion

- ✦ DUNE's ability to measure δ_{CP} depends on a precise understanding of $E_{true} \rightarrow E_{rec}$
 - ✦ Significant model dependence exists due to missing neutrons, missed or mis-ID pions, binding energy, etc.
- ✦ The DUNE-PRISM measurement program has been recommended by the DUNE ND group to provide a data-driven constraint of $E_{true} \rightarrow E_{rec}$
- ✦ The mechanism for making off-axis measurements (continuous vs discrete movement, and detector width) is in the process of detailed engineering studies
 - ✦ Decision still pending on whether downstream tracker will move, or whether a separate, downstream muon spectrometer is needed to move with the LAr
- ✦ Additional physics studies are underway to further demonstrate the robustness of this approach to neutrino interaction uncertainties